

# **Computer Setup, Git, and Python**

# Local/laptop computer Setup

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1. We will not be using Jupyter notebooks, so choose a text editor to use for writing code. I use MacVim, but Atom and Sublime Text are other popular options.
  2. Select a text editor and prepare it for class.
    - You may want to alias it, for example:  
csh: alias atom 'open -a atom'  
bash: alias atom='open -a atom'
  3. Install the VPN Pulse Secure software (see links page).
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# Python Setup

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1. Complete the Python Primer. Key things to note:

- Create the class conda environment.
- Always use `if __name__ == "__main__":`!
- Familiarize yourself with list of best practices and naming conventions.

**Stop here with pre-class preparation.**

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# Python Rec Arrays

# Rec arrays

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- Rec arrays thought of as single entities that can contain any number of variables (or arrays) by name. Similar to a Python structured array, but with different calling options.
  - rec arrays are very useful. They keep track of information (i.e. which column of a file contains the right ascension, which contains the declination)
  - They make reading files and sharing files extremely easy (*PyFITS* can read a file with millions of rows in a few seconds)
  - rec arrays are single objects. So, for instance, a function can return one entire rec array that contains a complex set of variables and arrays
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## Rec arrays

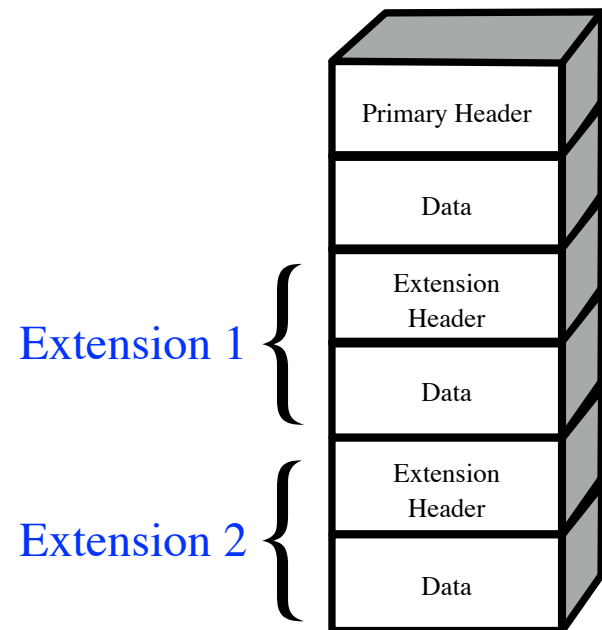
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- To learn how to make rec arrays and write them out as fits files, consult the documentation for *PyFITS* in *astropy*, linked from the syllabus, under week 1.

# FITS files

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- Moving forward, we will start to work with FITS files, which are a binary file format for storing rec arrays
- Although originally developed to transfer digital images FITS (Flexible Image Transport System) files are highly convenient for storing “tagged” information.
- They have “layers” of logical header/data units (HDUs) and are based on the concept of a record, or “rec” array



Schematic of a FITS file

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## The point of a rec array

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- I've put a rec array "struc.fits" in my week 1 Git directory and on the website. To read it using *PyFITS*:
    - *from astropy.io import fits*
    - *fx = fits.open(file)*
  - To see what the fits file contains try printing *fx.info()*
  - To access the data in the binary table, try *objs = fx[1].data* and to get its header *hdr = fx[1].header*
  - To use the variables (as you have used other arrays) you can try (after importing *matplotlib.pyplot* as *plt*)
    - *plt.plot(objs["RA"], objs["DEC"], "bx")*
    - *plt.show()*
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# Python tasks (Remember to commit to Git!!!)

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1. Write your code with proper structure (i.e., if `__name__ == "__main__"`).
  2. Read in my 'struc.fits' file and plot  $\delta$  vs.  $\alpha$  (Declination against Right Ascension) for objects in the file
  3. The *extinction* tag in 'struc.fits' is a 5-array. To access its first column you can use `objs["EXTINCTION"][:,0]`
  4. On your plot, overplot the  $(\alpha, \delta)$  of just those objects in 'struc.fits' where the first column of extinction is more than 0.22...the *numpy.where* function will be useful
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## Python tasks (Remember to commit to Git!!!)

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5. Generate 3 different sets of 100 random integers (see *numpy.random.randint*)
  6. Create a rec array with the tags *ra*, *dec*, and *randomnum* to store this information. Take *ra*, *dec* from *struc.fits*. Make *randomnum* a 3-array (see *numpy.reshape* if necessary). Write your rec array to a fits file.
  7. Experiment with docstrings versus comments:
    - *import polycalc*
    - *print(help(polycalc))*
    - *print(polycalc.get\_poly\_o3.\_\_doc\_\_)*
    - Remove *if \_\_name\_\_ == “\_\_main\_\_”* and *import*.
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