

# SQL Queries

# Modern astronomical surveys

---

- With the advent of the digital age, driven by the use of CCDs in cameras, astronomical surveys have started to become semi or fully automated
  - So, huge amounts of data are now arriving from sky surveys (Tb=Terabyte, Pb = Petabyte = 1000Tb)
    - ~50Tb of reduced data products over ~10 years of the Sloan Digital Sky Survey (SDSS)
    - ~2Pb of reduced data products over ~5 years of the Dark Energy Survey (DES)
    - ~50Pb of reduced data products over ~5 years of LSST (Large Synoptic Survey Telescope) operations
-

# Mining modern astronomical surveys

---

- With such a large amount of data to sift through, astronomers have become more involved in developing data mining techniques
  - We've discussed aspects of this in terms of pixelating the sky...which is really a method for indexing large amounts of data in a database for efficient searches
  - The HTM index, a type of quad-tree that we've discussed briefly, is an efficient schema for storing data and searching through that data by object position
  - We won't discuss the math of HTM in detail (a good description is linked from the syllabus) but think of it as a HEALPix-like index, coupled with the spherical cap formalism to find which HTM pixels lie in a cap
-

# Introduction to SQL

---

- Visit the *SDSS SQL Tutorial* linked from the syllabus
  - Read and/or try the following tutorials:
    - 1. *Introduction*
    - 2. *A simple Query*
    - 3. *Common Searches*
  - Note though, that nothing in these first 3 SQL tutorials makes use of the HTM indexing scheme
  - The genius of HTM is coded in functions such as, e.g., *fGetNearbyObjEq( $\alpha, \delta, \theta$ )* which can *very* rapidly find objects at a radius  $\theta$  around a position ( $\alpha$ =RA,  $\delta$ =dec)
    - Try *SDSS SQL Tutorial 10. Functions*
-

# Python tasks

---

1. Using the *SDSS SQL Search Box* (see the link from the syllabus) download the RA, Dec and g-band magnitude for *all* objects in the SDSS that are within a radius of 2' of the position  $(\alpha, \delta) = (300^\circ, -1^\circ)$ 
    - You should recover about 350 total objects
    - Make sure to return *all* objects, not just objects of a specific *type*
    - *type=3* corresponds to “galaxies” or, more precisely, “objects that are resolved and extended in imaging”
    - *type=6* corresponds to “stars” or, “objects that are unresolved and appear as point sources in imaging”
  2. Write Python code that reads in these SDSS objects and plots RA against dec. *Use circles for your data points*
-

# Python tasks

---

3. Repeat your plot, but bin your points such that objects with *larger*  $g$  are plotted using *smaller* circles (i.e. plot  $16 < g < 17$  at a larger size than  $17 < g < 18$ )
    - Using Matplotlib, `plt.scatter(ra,dec,s=s)` will allow you to plot points of different sizes
    - Here,  $s$  is the “size” of the point, but note that it’s actually the *area* of the marker (so multiplying  $s$  by 4 will double the radius of the plotted point)
  4. Use the *SDSS Navigator Tool* linked from the syllabus, to display an image near  $(\alpha, \delta) = (300^\circ, -1^\circ)$ . Zoom in until the image is  $\sim 2'$  across (the scale will be  $\sim 20''$ )
    - Check that your plot looks reasonably like the *SDSS Navigator Tool* image
-

# Python tasks

---

5. Write a function or class that executes SQL queries directly from Python. See links in the syllabus for ideas.
6. Execute the same tasks, but directly from Python.