SQL Queries

Modern astronomical surveys

- With the advent of the digital age, driven by the use of CCDs in cameras, astronomical surveys have started to become semi or fully automated
- So, huge amounts of data are now arriving from sky surveys (Tb=Terabyte, Pb = Petabyte = 1000Tb)
 - ~50Tb of reduced data products over ~10 years of the Sloan Digital Sky Survey (SDSS)
 - ~2Pb of reduced data products over ~5 years of the Dark Energy Survey (DES)
 - ~60Pb of reduced data products over 10 years of LSST (Legacy Survey of Space and Time) operations

Mining modern astronomical surveys

- With such a large amount of data to sift through, astronomers have become more involved in developing data mining techniques
- We've discussed aspects of this in terms of pixelating the sky...which is really a method for indexing large amounts of data in a database for efficient searches
- The HTM index, a type of quad-tree that we've discussed briefly, is an efficient schema for storing data and searching through that data by object position
- We won't discuss the math of HTM in detail (a good description is linked from the syllabus) but think of it as a HEALPix-like index, coupled with the spherical cap formalism to find which HTM pixels lie in a cap

Introduction to SQL

- Visit the SDSS SQL Tutorial linked from the syllabus
- Read and/or try the following tutorials:
 - 1. Introduction
 - 2. A simple Query
 - 3. Common Searches
- Note though, that nothing in these first 3 SQL tutorials makes use of the HTM indexing scheme
- The genius of HTM is coded in functions such as, e.g., $fGetNearbyObjEq(\alpha, \delta, \theta)$ which can *very* rapidly find objects at a radius θ around a position (α =RA, δ =dec)
 - Try SDSS SQL Tutorial 10. Functions

Python tasks

- 1. Using the SDSS SQL Search Box (see the link from the syllabus) download the RA, Dec and g-band magnitude for *all* objects in the SDSS that are within a radius of 2' of the position $(\alpha, \delta) = (300^{\circ}, -1^{\circ})$
 - You should recover about 350 total objects
 - Make sure to return *all* objects, not just objects of a specific *type*
 - *type=3* corresponds to "galaxies" or, more precisely, "objects that are resolved and extended in imaging"
 - *type=6* corresponds to "stars" or, "objects that are unresolved and appear as point sources in imaging"
- 2. Write Python code that reads in these SDSS objects and plots RA against dec. *Use circles for your data points*

Python tasks

- 3.Repeat your plot, but bin your points such that objects with *larger* g are plotted using *smaller* circles (i.e. plot 16 < g < 17 at a larger size than 17 < g < 18)
 - Using Matplotlib, *plt.scatter(ra,dec,s=s)* will allow you to plot points of different sizes
 - Here, *s* is the "size" of the point, but note that it's actually the *area* of the marker (so multiplying *s* by 4 will double the radius of the plotted point)
- 4. Use the *SDSS Navigator Tool* linked from the syllabus, to display an image near $(\alpha, \delta) = (300^{\circ}, -1^{\circ})$. Zoom in until the image is ~ 2' across (the scale will be ~ 20'')
 - Check that your plot looks reasonably like the *SDSS Navigator Tool* image

Python tasks

- 5. If you feel like a challenge (i.e. you will not need to have written this function for class), write a function or class that executes SQL queries directly from Python. See links in the syllabus for ideas.
- 6. Execute the same tasks, but directly from Python.